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LIST OF NEBULAS OBSERVED AT THE LEANDER McCORMICK  
OBSERVATORY, AND SUPPOSED TO BE NEW

No.	R.A. 1890.0	Decl. 1890.0	Mag.	Size	Form	Condensation	No. of Obs.	Obsvr	Notes
1	0 10 <sup>h m</sup>	-22° 3'	14.0	pS	iR, 1E 120°	gbM	1	S	
2	0 10	14 7	14.0	0'.7	R		1	L	
3	0 19	14 35	15.5	0.2	R		1	L	
4	0 20	5 46	pF	vS	R	lbM	1	L	
5	0 25	13 58	14.5	0.4	R		1	L	
6	0 28	11 22	13.0	S	R	bsp	1	M	*12 P 90° Δ 3'.2
7	0 32	14 47	F	S	mE 0°	bM	1	S	Faint wing sp
8	0 33	15 16	F	S	mE 150°	bM	1	S	
9	0 34	14 51	eF	eS	E 60°	gbsbMN	1	S	
10	0 35	-19 14	12.0	vS	R	glmbMN	2	L	
11	0 35	-14 24	15.5	0.8	pE 0°	lbMN	1	L	G.C. 107?
12	0 39	4 23	pF	pS	E 20°	bMN	1	L	*8, f 20 <sup>s</sup>
13	0 43	13 44	16.0	0.2	1E 30°		1	L	
14	0 47	13 46	16.0	0.5	R		1	L	1st of 4
15	0 47	13 46	16.0	0.5	R		1	L	2nd of 4
16	0 47	13 46	16.0	0.5	R		1	L	3d of 4
17	0 47	13 44	16.0	0.5	R		1	L	4th of 4
18	0 50	11 16	15.5	S	iR	gbM	1	M	*8, p 30 <sup>s</sup>
19	0 50	11 15	16.0	vS			1	M	Neb? f (18), P 75° Δ 1'.0
20	0 50	-17 16	16.0	vS			1	L	F*, p 40 <sup>s</sup>
21	0 54	-18 52	14.0	pS	pE	bnpN	2	L	
22	0 54	19 1	14.0	vS	R	sbMN	1	L	
23	0 59	17 9	16.0	eS	R		1	L	
24	1 0	18 24	vF	vS	R	gbM	1	L	
25	1 1	20 38	14.0	vS	vE	sbMN	1	L	
26	1 5	20 36	13.0	eS	R	sbMN	1	L	Neb?
27	1 16	16 57	15.0	vS	R		1	L	
28	1 20	18 46	13.2	vS	R		2	L	
29	1 20	20 32	13.0	vS	R	sbMN	1	L	
30	1 22	-18 46	14.0	pS	1E 70°	bMN	2	L	sev F st f in line n and s
31	1 23±	-10 51	15.0	vS	R		1	L	
32	1 27	17 7	12.5	pS	E 225°	glbM	1	L	
33	1 33	20 30	eF	vS	R		1	L	
34	1 35	18 25	vF	vS	v1E	sbMN	1	L	
35	1 36	13 36	15.5	0.1		gbMN	1	S	
36	1 44	27 58	14.0	1.3	R	gbM	1	S	
37	1 44	17 17	14.0	vS	R	lbMN	1	L	
38	1 47	13 22	16.0	0.4		gbMN	1	S	
39	1 49	17 5	vF	vS	R		1	L	
40	1 49	-11 21	14.5	pL	iR		1	M	*9, P 90° Δ 3'.6

No.	R. A. 1890	Decl. 1890	Mag.	Size	Form	Condensation	No. of Obs.	Obsv'r	Notes
41	<sup>h</sup> 1 50 <sup>m</sup>	—17° 34'	14.0	vS	R	bMN	2	L	*11, p 11 <sup>s</sup>
42	1 51	17 16	12.0	vS	R	bMN	2	L	env 14.0
43	1 51	9 27	11.5	0.4		gbMN	1	S	
44	1 52	9 31	13.0	1.6×0.4	E 65°	gbMN	1	S	
45	1 52	16 34	16.0	0.2	iR	gbM	1	S	
46	1 54	9 26	14.0	0.2	R	gbMN	1	S	
47	2 1	16 17	14.8	0.3	R	gbM	1	S	
48	2 1	16 21	15.8	0.1	R	gbM	1	S	
49	2 6	10 51	15.5	eS	E 325°	gbM	1	S	*10, n 1'.0
50	2 11	—18 19	vF	pS	vE 0°	gvlbM	3	L	sev vF st inv; n end b?
51	2 17	—17 12	15.5	vS	R?		2	L	
52	2 20	16 20	15.0	0.4	R	gbM	1	S	
53	2 23	11 19	13.0		R		1	M	} neb* *?
54	2 23	11 18	13.5		R		1	M	
55	2 23	15 1	15.0	0.4	vE 0°	sbN like a * *	1	L	
56	2 23	10 59	15.0	0.7	R	gbM	1	S	
57	2 24	11 31	15.0	0.4		gbM	1	S	
58	2 27	19 8	14.5	S		gbM	1	S	
59	2 28	11 27	15.0	0.4		gbM	1	S	
60	2 30	—17 0	13.0	vS	R	bMN	3	L	
61	2 34	—11 30	15.0	0.4			1	S	neb? in same field with G.C. 5262 and 5263
62	2 34	16 17	13.5	vS	IE 0°?	lbM	2	L	2 B st, p 20 <sup>s</sup>
63	2 38	16 46	15.5	vS	R		1	L	
64	2 38	16 41	14.0	vS		bMN	1	L	
65	2 41	18 1	13.0	vS	R	sbMN	1	L	
66	2 41	18 1	13.5	vS	R	sbMN	1	L	
67	2 42	18 9	12.0	vS	R	glsbMN	1	L	1st of 3
68	2 42	18 11	13.0	pS	IE 0°	slbMN	1	L	2nd of 3
69	2 42	18 10	12.5	vS	IE 30°	sbMN	1	L	3d of 3
70	2 42	—16 35					1	L	same as (64)?
71	2 43	—16 10	14.5	vS	R		1	L	
72	2 45	18 29	12.0	eS	R		1	L	*?
73	2 45	14 55	14.0	0.5	R	gbsbMN	1	L	env 14.5
74	2 46	26 10	16.0	eeS	iR	gbM	1	S	*9, nf 1'.0
75	2 49	14 58	14.0	0.3	R	gbMN	1	L	env 15.0
76	2 52	15 29	14.0	0.4	R	sbMN	1	L	env 15.0
77	2 52	15 27	16.0	0.4	R		1	L	neb? in same field with (76)
78	2 53	15 34	15.0	0.7	E 0°	sbMN	1	L	env 16.0
79	2 53	14 49	15.0	0.4	R	sbMN	1	L	env 15.5
80	2 54	—17 37	vF	pS	vE 75°	bnp	1	L	spindle shaped
81	2 57	—19 20	15.5	pS		gbM	1	S	*12 f 1'.0
82	2 57	15 27	15.0	vS	R	bMN	1	L	env 16.0
83	2 57	15 29	15.0	vS	R	bMN	1	L	env 16.0
84	2 57	10 6	15.5	0.7×0.3	E 120°		1	S	*10, P 240° Δ 2'.5
85	3 0	14 48	14.0	0.3	R	bMN	1	L	neb? env 14.5
86	3 0	12 46	15.5		E 45°		1	L	B* and sev F st inv in neb, r
87	3 0	10 7	14.0	0.7×0.3	E 25°		1	S	*9.5, P 240° Δ 3'.0
88	3 1	26 10	15.0	vS	iR, E 340°?	gbMN	1	S	
89	3 1	15 56	15.0	vS	R		1	L	
90	3 1	—16 2	15.5	vS	R		1	L	1st of 5
91	3 1	—16 5	15.5	vS	R		1	L	2nd of 5
92	3 1	16 7	15.5	vS	R		1	L	3d of 5
93	3 1	16 6	15.5	vS	R		1	L	4th of 5, 5th is G.C. 648
94	3 2	9 58	14.0	0.7×0.2	E 60°		1	S	48s = f G.C. 647 same decl. stell N in cen of vF neb; 1st of 3; *10, P 15° Δ 3'.0.
95	3 2	10 3	15.5	0.4		dif	1	S	2nd of 3
96	3 2	10 2	14.5	0.2		stell	1	S	3d of 3
97	3 5±	—16 0	16.0	pL	E like fan		1	L	neb?

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No.	R. A. 1890.	Decl. 1890	Mag.	Size	Form	Condensation	No. of Obs.	Obsv'r	Notes
98	3 <sup>h</sup> 8 <sup>m</sup>	-26° 10'	15.5	4.1×2.0	E 315°	vgbsbMN	1	S	
99	3 10	16 18	15.0	pS	iR	sbMN	1	L	env 15.5
100	3 10	-15 31	14.0	0.7	IE 0°	sbM	1	L	env 16.0
101	3 13	-14 23	15.8	0.1			1	S	near (102)
102	3 14	14 24	15.0	0.2		gbMN	1	S	*10, P 75° Δ 3'.0
103	3 15	18 57	13.0		iF, vmE 135°		1	S	
104	3 16	25 56	14.8	vS		gbM	1	S	no N, * 10.5, f 4'.0
105	3 20	26 4	16.3	vS			1	S	neb?
106	3 30	25 18	13.0				1	S	*9, nf 5'.0
107	3 30	16 4	14.0	vS	R	lbM	1	L	
108	3 31	16 16	14.0	vS	R	glbM	1	L	
109	3 33	16 17	14.5	vS	R	lbM	1	L	
110	3 35	-15 53	16.0	pL	vE 150°	glbM	1	L	sev vF st inv
111	3 36	-15 57	15.0	vS	R	lbM	1	L	
112	3 40	23 21	15.0		vmE 60°	N	1	S	*10, f 1'.0
113	3 40	9 33	16.0	0.8		stell N	1	S	} D, P 310° Δ 0'.4
114	3 40	9 33	16.0	0.8		stell N	1	S	
115	3 42	25 52	15.0	pS		gbM	1	S	
116	3 48	15 44	14.5	vS	pE 45°		1	L	
117	3 48	8 54	14.0	0.1		stell N	1	S	1st of 3
118	3 49	8 54	15.0	0.2			1	S	2nd of 3
119	3 49	8 52	15.0	0.2			1	S	3d of 3
120	3 57	- 9 38	16.0	0.2	R	gbM	1	S	
121	3 57	- 9 37	15.3	0.3	R	gbM	1	S	
122	3 59	11 28	14.0	0.1	R	gbM	2	S	
123	4 5	9 5	16.0	0.7	R	gbMN	1	S	} D, P 340° Δ 0'.5
124	4 5	9 5	16.0	0.7	R	gbMN	1	S	
125	4 10	13 31	16.0	0.1	R	gbM	1	S	
126	4 13	18 9	pF	pS	iR		1	L	cl? or neb with sev vF st and one * 11.5 n of cen
127	4 18	16 7	14.0	vS	IE 170°	glbM	1	L	inv *8, p 6°
128	4 18	16 2	14.5	eS	R	glbM	1	L	
129	4 18	16 0	15.5	vS	R	lbM	1	L	} D
130	4 18	-16 0	15.0	vS	R	lbM	1	L	
131	4 18	-16 2	16.0	pS	IE		1	L	
132	4 24	17 52	12.5	vS	R	sbMN	1	L	env 13.0
133	4 24	17 47	12.5	eS	R	sbMN	1	L	env 14.0
134	4 26	11 31	15.5	0.2	R	gbM	2	S	
135	4 31	13 46	15.8	0.2	R	gbM	1	S	
136	4 36	20 38	14.0	0.4	E 40°	gbMN	1	S	
137	4 40	16 6	13.0	pS	E 0°	glbMN	1	L	env 14.0
138	4 47	15 32	15.5	vS	vE 30°		1	L	
139	4 49	20 44	15.0	0.2	R		1	S	
140	4 53	-15 28	13.0	vS	R	glbMN	2	L	env 13.5
141	4 55	-16 0	12.0	vS		sbMN	2	L	} D, P 110° Δ 10"
142	4 55	16 0	12.0	vS			2	L	
143	4 56	18 19	14.0	0.4	E 45°		1	S	
144	4 56	18 19	15.0	0.2	E? 90°		1	S	in same field with (143)
145	5 0	19 36	15.0	0.1		gbM	1	S	
146	5 2	18 19	14.5	0.1		gbM	1	S	
147	5 6	15 15	14.0	vS	IE 140°		1	L	
148	5 20	16 4	15.0	pS	pE 0°	glbM	1	L	
149	5 28	14 12	15.0	pS	R	bmN	1	L	env 15.5
150	6 14	-18 29	14.0	pS	E 45°	glbMN	1	L	
151	6 14	-18 29	15.5	vS	R		1	L	in same field with (150)
152	8 43	13 53	16.0	0.4			1	S	neb?
153	8 49	- 4 34	15.8	0.2		stell	1	S	

No.	R.A. 1890	Decl. 1890	Mag.	Size	Form	Condensation	No. of Obs.	Obs'r	Notes
154	<sup>h</sup> 9 <sup>m</sup> 22	-14° 3'	15.0	0.3	R	bMN	1	L	env 15.5
155	9 26±	+31 36±					1	S	found while looking for Winnecke's comet
156	9 33	-11 56	15.0	1.1	iR	gbM	1	L	
157	9 37	16 9	eeF	S	R	gbM	1	S	
158	9 43	5 59	eF	vS	R		1	L	
159	9 50	11 57	15.5	1.2	D or bi N		1	L	tri N?
160	9 51	-25 11	14.5	2.7×0.8	E 120°		1	S	16.5 vgb 16.0 vsb MN
161	9 51	- 5 51	16.0	vS	E 45°		1	L	
162	9 54	5 54					1	L	same as (161)?
163	9 59	20 15	15.0	0.1	R		1	S	neb?
164	10 2	15 36	16.0	1.6×0.7	vE 45°		1	L	
165	10 2	15 36	16.0	1.3×0.7	vE 170°	lbN	1	L	
166	10 4	16 6	15.5	0.5	R	sbMN	1	L	
167	10 4	16 6	16.0	0.3	R		1	L	
168	10 6	20 20	15.0	0.4	R	gbM	1	S	
169	10 9	20 5	15.5	0.4	R	gbM	1	S	
170	10 10	-18 29	16.0	0.8	iR		1	S	neb?
171	10 14	-25 16	15.8	1.6	iR	gbM	1	S	
172	10 18	21 42	16.0	1.6	spiral?	stell N	1	S	N 0'.1, env 16.0
173	10 30	12 5	15.0	1.6			1	L	D neb or sev st inv in neb
174	10 30	12 9	15.0	0.7	R	gpmbM	1	L	
175	10 30	12 7	15.5	0.5	iR	gbMN	1	L	
176	10 32	24 45	15.5	0.1	R	gbM	1	S	*15, n 3".0
177	10 41	24 40	16.0	0.2	R		1	S	
178	10 50	20 30	15.0	1.3×0.4	E 125°		1	S	
179	10 52	19 3	15.5	0.4	R	gbM	1	S	
180	10 53	-14 22	15.5	0.6	E 90°	gbMN	2	S	
181	10 57	-14 22	16.0	1.2	iR	glbM	1	L	
182	10 59	5 58	15.0	0.1	1E 110°		1	S	in same field with neb disc by Stephan
183	11 1	18 52	12.0	0.8		gbMN	1	S	
184	11 4	17 41	14.0	2.5×0.4	E 95°	gbsbgbM	1	S	G.C. 2330?
185	11 5±	19 27	vF	vS	R	gbMN	1	S	1st of 2
186	11 5±	19 27	eeF	eS	R	gbM	1	S	2nd of 2
187	11 24	8 1	15.0	0.2		gbMN	1	S	
188	11 25	2 37	14.5	0.3	R	gbM	1	S	
189	11 25	2 39	16.0	0.1			1	S	
190	11 33	- 2 48	15.0	0.2		gbM	1	S	
191	11 33	-11 57	15.0	0.8	iR	glbM	1	L	S* or neb f
192	11 44	9 7	15.5	1.8	iR	gbM	1	S	
193	12 8	11 58	15.8	0.8			1	S	
194	12 54	13 32	16.0	0.3	R		1	L	
195	12 54	13 27	15.5	0.3	vE 45°	sbMN	1	L	
196	13 8	18 58	14.0	1.6×0.2	E 30°	gbM	1	S	8' f G.C. 3448 P 110°
197	13 13	9 38	15.0	0.2			1	S	
198	13 19	9 37	14.0	0.2		gbM	1	S	
199	13 53	13 37	12.0	S	R	vgbM	1	M	
200	13 57	-14 4	F	pS	vE	gbp	1	L	a little curved, shades off gradually like a comet's tail; no N seen
201	14 8	-17 27	14.0	0.4	R	gbM	1	S	*13 in field
202	14 14	18 37	14.0	0.7	iR	gbMN	1	S	
203	14 27	16 6	16.0	0.2	R	glbM	1	L	
204	14 27	14 6	pF	S	E	gbM	1	L	
205	14 36	17 58	12.8	0.3	R	gbM	2	S	*10.5, np 2'.7
206	14 38	20 25	vF	S	1E	glbM	1	L	
207	14 39	11 28	14.0	vS	R	sbMN	1	L	
208	14 39	11 22	12.0	pS	pmE	gbMN	1	L	env 15.0
209	14 39	20 25	F	S	vE	smbMN	1	L	
210	14 39	-18 1	15.5	0.2			1	S	neb?

No.	R.A. 1890	Decl. 1890	Mag.	Size	Form	Condensation	No. of Obs.	Obsv'r	Notes
211	14 43 <sup>h m</sup>	-19 49	14.5	0.3	R	glbMN	1	L	
212	14 46	20 55	eF	1.0	R	gbM	1	S	
213	14 53	16 14	15.4	0.8×0.3	IE 135°	bMN	2	L	
214	14 53	16 10	12.8	0.8		gbsmbMN	2	L	N almost stell
215	14 54	13 25	vF	vS		sbM	1	L	1st of 3
216	14 54	13 26	vF	vS		sbM	1	L	2nd of 3 brightest and most nebulous of the three
217	14 54	13 24	vF	vS		sbM	1	L	3d of 3
218	14 55	17 25	eF	vS	IE 230°		1	S	bet 2 vFst
219	14 56	16 23	15.0	0.8	E 10°		1	L	F*? inv in neb * on p side of neb
220	14 56	-15 42	11.0	0.8		gbMN	1	S	stell
221	14 56	-15 46	14.0	0.8			1	S	sev F st in field
222	15 1	14 8					1	L	*13.0 inv in vF neb 3st, p 14.0; *8.0 f 10.0 15.0 s
223	15 4	18 0		S	R	gbMN	1	S	*12 in eF neb
224	15 8	14 8	15.2	0.2	R	lbsbMN	2	L	in field with Harv. 331
225	15 10	17 9	vF	vS	E 235°		1	S	
226	15 10	11 7	13.0	pS	IE	gbM	1	L	*11, f
227	15 10	14 35	16.0	2.5		gbM	1	S	
228	15 31	16 13	14.5	pS	vIE?	glsbMN	2	S	
229	15 36	12 52	eF	vS		sbMN	1	L	in a group of st
230	18 2	-29 34	vF	vS	R		1	L	rr
231	20 26	-25 51	13.2	pS	R	sbMN	2	L	* n 14.0, neb * in field?; env 14.0
232	20 28	11 46	15.0	0.6×0.2	E 120°	gbM	1	L	divided into 2 parts?
233	20 30	25 40	13.8	vS	R	slbMN	2	L	env 14.0
234	20 50	11 30	16.0	0.2	iR		1	L	
235	20 50	19 0	14.0	vS	R	glbMN	2	L	
236	20 53	26 7	14.0	vS	R	sbMN	1	L	env 15.0
237	21 1	25 56	14.0	eS	R	bMN	2	L	1st of 3
238	21 1	25 56	15.0	vS	R	bMN	2	L	2nd of 3
239	21 1	25 54	vF	vS	vIE	glbM	2	L	3d and b of 3
240	21 1±	-24 51	15.0	vS		sbMN	1	L	
241	21 4	-20 57	14.0	vS	iR	sbnMN	1	L	
242	21 35±	10 50	11.0	vS	E 310°?	smbMN	1	L	*n, P 310°
243	21 38	25 52	vF	pS	vE 90°		3	L	like comet with tail; 2 st inv
244	21 50	25 54	14.0	vS	R	sbMN	1	L	B* *, p 8°; env 16.0
245	21 55	5 59	13.0	pS	iR	bMN	2	L	*p 36°
246	21 57	33 19	pF	pS	R	lbM	1	L	
247	21 57	20 52	eF	pS	E	lbM	1	L	
248	22 12	24 17	pF	vS	R	gvlbM	1	L	B* * p 13°, F* *, f 5°
249	22 16	16 7	16.0	2.5×0.5	vE 30°	sbMN	1	L	
250	22 19	-13 46	eF	pS	R	glbM	1	L	
251	22 24	-25 59	vF	vS	R		1	L	
252	22 26	18 9	vF	pS	IE 0°	lbM	1	L	
253	22 28	13 33	pB	vS	R		1	L	no * in field
254	22 29±	23 5	15.0	pS	R	bMN	1	L	
255	22 33±	23 16	pF	pS	E	lbM	1	L	
256	22 38	24 17	pF	vS	pmE	bMN	1	L	
257	22 39	20 32	14.5	eS	R	gbMN	2	L	*11, nf 4'.0; neb *
258	22 43	20 20	eF	vS	R	gbM	1	L	
259	22 51	11 32	15.5		IE? 90°		1	M	*10, P 260° Δ 4'.0
260	22 54	- 7 38?	14.0	0.8	iR		1	S	*10, p
261	23 0	-20 32	14.0	vS	R	gbM	1	S	
262	23 13±	7 31	13.5	pS	IE 180°	lbMN	1	L	
263	23 20	19 41	vF	vS	R	bMN	1	L	
264	23 33	23 36	eF	vS	R		1	L	
265	23 37	20 4	15.0	eS		gbM	1	S	bet 2 st 12
266	23 42	17 14	16.0	vS			1	L	
267	23 43	-17 10	14.0	vS	R		1	L	

No.	R.A. 1890	Decl. 1890	Mag.	Size	Form	Condensation	No. of Obs.	Obsvr	Notes
268	23 <sup>h</sup> 44 <sup>m</sup>	-14° 0'	13.0	vS	R	gbM	1	S	* 10, p 8'.0
269	23 45	17 13	15.0	vS	R		1	L	F * f
270	23 47	-14 0±	15.0	vS	IE 315°	gbM	1	S	near (268)
271	23 56	-19 23	16.0	pS	iF E 90°?		1	S	
272	23 56	11 21	15.0	vS	R	stell N	1	M	* 8.5, P 240° Δ 3'.6
273	23 59	-17 6	14.0	pS	iF E 125°	glbM	2	S	

S = Ormond Stone; L = F. P. Leavenworth; M = Frank Muller; env = envelope; P = position angle; Δ = distance; other abbreviations as in *Herschel's general catalogue*.

In the earlier observations *Herschel's abbreviations* were used to designate brightness and size. Afterwards numerical magnitudes were employed to indicate brightness, assuming that the faintest nebula visible in the 66<sup>cm</sup> refractor, with power 167, is 16.3, that being the theoretical limit for *University of Virginia*, 1886 October 12,

stars. The magnitudes given refer to the nucleus, or, in case there is no nucleus, to the brightest part. Still later the custom was instituted of estimating the diameters of the nebulas observed in fractions of the diameter of the field, and from these deducing their dimensions in minutes of arc.

Sketches have been made of the larger portion of the nebulas contained in this list.

ORMOND STONE.

## LATITUDE OF THE SAYRE OBSERVATORY.

BY PROF. C. L. DOOLITTLE.

During the years 1876-77 I made a somewhat extended series of observations with the Zenith-Telescope for determining the latitude of our observatory. A few observations were also made in the spring of 1878; the mean date of the series being 1877.17.

Sixty pairs of stars were employed; the number of observations was 459, an average of 7.65 per pair.

Sixty-two of the stars were taken from Boss's Catalogue of 500 stars, the remaining 58 were reduced by myself, nearly all the data to be found in the library of the U. S. Naval Observatory being employed. Boss's systematic corrections were applied in order to make the entire series of declinations homogeneous. The results of this work were published in the *Astronomische Nachrichten*, No. 2260.

In September, 1885, I began a series of observations for redetermining the latitude, and the observations were finished August, 1886. The mean date was 1886.12. The stars employed were the same as in the previous determination, with the exception of three pairs, which for different reasons were not available. This series comprised 288 observations of 57 pairs, an average of 5.1 per pair.

The observations were all made by myself, the same instrument being used throughout. The level-tube was twice refilled during the interval between the two series; otherwise nothing was done by way of change or repair. As the instrument was used more or less every year by the students of the University, there must have been some wear of the parts.

The precision of the observations comprising the latter series was however a little greater than of the former, therefore no great deterioration could have taken place.

Taking the results of the earlier series, and using only those stars which were employed in the latter, 57 equations were written, of the form

$$\Delta\varphi - \frac{1}{2}(M - M') \Delta R = \varphi - \varphi'$$

each pair of stars giving one equation.

$\Delta\varphi$  is the correction to the assumed latitude  $\varphi'$ .

$\Delta R$ , the correction to the assumed value of the micrometer-screw.

The results were as follows:

$$\begin{aligned}\varphi_1' &= 40^\circ 36' 23.887 \\ \Delta\varphi_1 &= +.018 \pm .037 \\ \varphi_1 &= 40 36 23.905 \pm .037.\end{aligned}$$

In the same manner the second series gave the following results:

$$\begin{aligned}\varphi_2' &= 40^\circ 36' 23.530 \\ \Delta\varphi_2 &= -.018 \pm .051 \\ \varphi_2 &= 40 36 23.512 \pm .051.\end{aligned}$$

Therefore  $\varphi_1 - \varphi_2 = 0''.393 \pm 0.063$ .

If we regard only those parts of the probable errors of  $\varphi_1$  and  $\varphi_2$  which are independent of the declinations of the stars employed, we find for the probable error of the difference, viz:  $(\varphi_1 - \varphi_2)$ ,  $r = .045$ .

The following tabular statement exhibits the values of the latitude, as given by each pair of stars employed. If we